

Introduction to the PTetra workshop

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Outline

Orbital motion-limited (OML) theory

Case study

From beginning to end

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Orbital motion-limited (OML) theory

Energy and momentum conserving orbits.

Current due to attracted species s :

$$I_s = C \underbrace{\overset{\text{area}}{A} q_s n_s \sqrt{\frac{kT_s}{2\pi m_s}}}_{I_{\text{th},s}} \underbrace{\left(1 - \frac{q_s V}{kT_s}\right)}_{\eta_s}^\beta$$

Current due to repelled species s :

$$I_s = I_{\text{th},s} \exp\left(\frac{q_s V}{kT_s}\right)$$

Mott-smith and Langmuir, DOI: 10.1103/PhysRev.28.727

	Plane	Cylinder	Sphere
C	1	$2/\sqrt{\pi}$	1
β	0	0.5	1

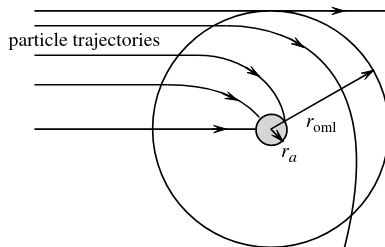


Figure: Orbital motions (From Bilén, PhD thesis, 1998)

Orbital motion-limited (OML) theory

Energy and momentum conserving orbits.

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Assumptions to OML theory:

- ▶ Radius $r \ll \lambda_{De}$
- ▶ Length $l \gg \lambda_{De}$ (for cylinders)
- ▶ $\eta_s > 2$ (for cylinders)
- ▶ No drift velocity
- ▶ No magnetic field

Orbital motion-limited (OML) theory

Langmuir probes:

Vary V , measure I , infer e.g. n_e .

E.g. Jacobsen *et al.*, DOI: 10.1088/0957-0233/21/8/085902

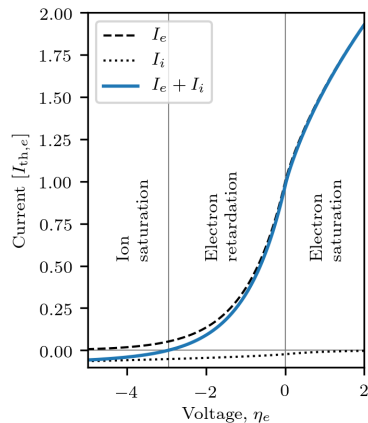


Figure: OML IV-characteristic for cylinder. (From Marholm, PhD thesis)

Orbital motion-limited (OML) theory

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Unconnected objects (spacecraft, dust, etc.)
in plasma settle at the *floating potential*:

$$V = \text{const} \Rightarrow I_e + I_i = C \frac{dV}{dt} = 0$$

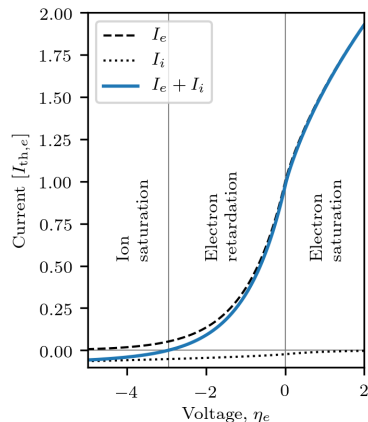


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Spaceborn probes are biased wrt. the
spacecraft, and the probe current also charge
the spacecraft!

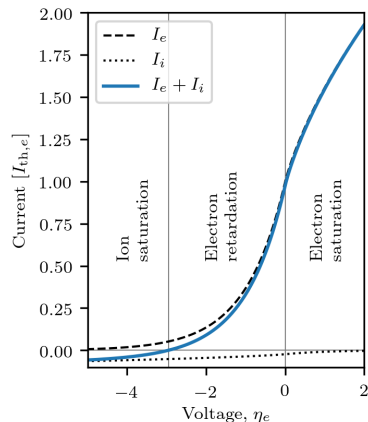


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- ▶ Each team given an object
 - ▶ Simulate the object
 - ▶ at 1V
 - ▶ at 2V
 - ▶ at 3V
 - ▶ split at 1V and 3V
 - ▶ at the floating potential
- at *ionospheric conditions*

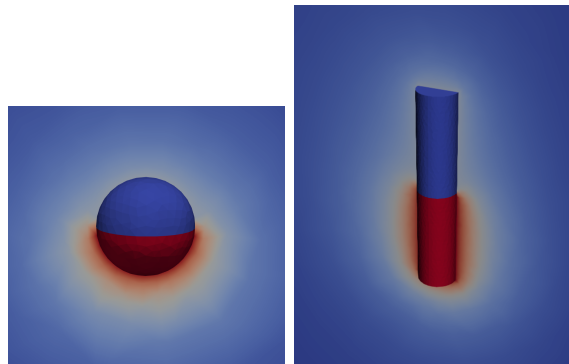


Figure: A spherical and cylindrical object where the two halves are biased at different voltages, along with the surrounding electric potential.

Ionospheric conditions

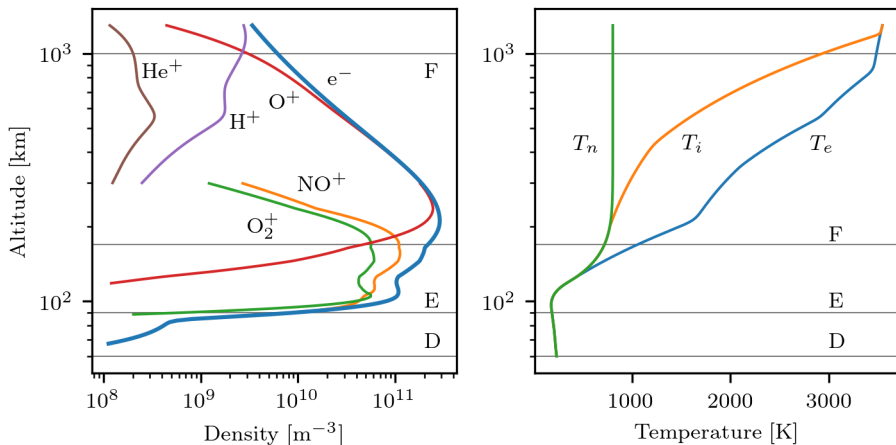


Figure: IRI-2016 model of the ionosphere, 45°N local noon. (From Marholm, PhD thesis)

Ionospheric conditions

Parameters for low Earth orbit simulations:

- ▶ $n_e = n_i = 10^{11} \text{ m}^{-3}$
- ▶ $T_e = T_i = 1000 \text{ K} = 0.0862 \text{ eV}$
- ▶ $m_i = 16 \text{ amu}$ (O^+)
- ▶ $u = 7000 \text{ m/s}$ (drift speed)

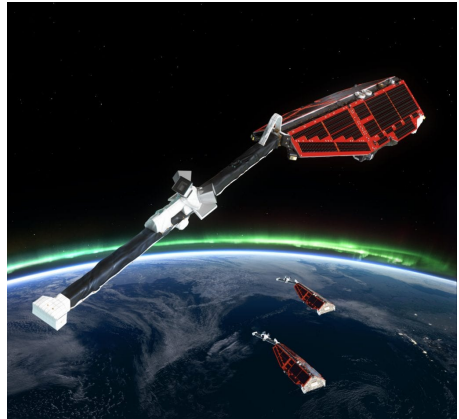


Figure: The Swarm satellites ESA/AOES

Medialab

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- ▶ $m_i = \cancel{16 \text{ amu}} \rightarrow 1/16 \text{ amu} \gg m_e$ (speed up sim.)
- ▶ $u = 7000 \text{ m/s}$ (drift speed)

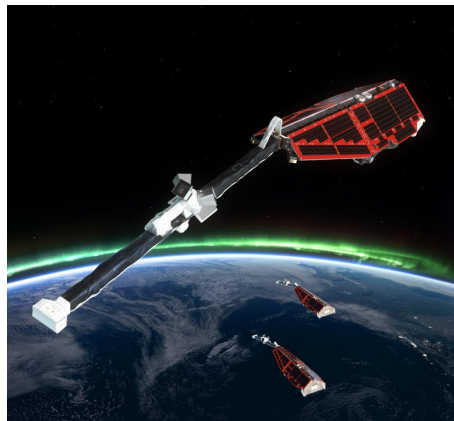


Figure: The Swarm satellites ESA/AOES

Medialab

Reduced mass ratio

Reduced m_i/m_e ...

- ▶ Speeds up simulations ($\tau \propto \sqrt{m_i}$)
- ▶ Changes the physics, e.g.:
 - ▶ increases ion current (see OML)
 - ▶ the mach cone half-angle

$$\theta = \arcsin\left(\frac{1}{M}\right), \quad M = \frac{u}{c_s},$$

where c_s is sound speed

$$c_s = \sqrt{\frac{kT_e + 3kT_i}{m_i}}.$$

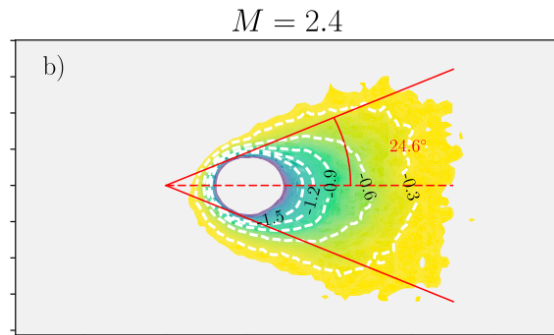


Figure: Mach cone superimposed on numerical simulation of rocket wake. (From Darian *et al.*, DOI: 10.1002/2017JA024284)

Ionospheric conditions

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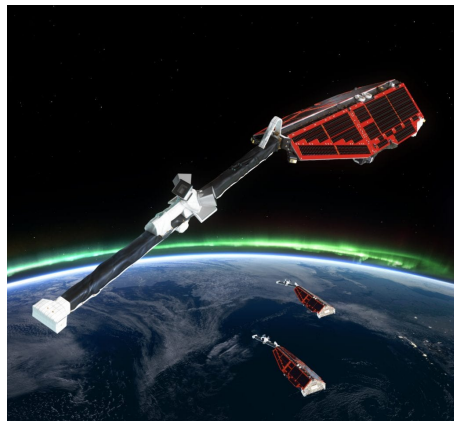


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Ionospheric conditions

Parameters for low Earth orbit simulations:

- ▶ $n_e = n_i = 10^{11} \text{ m}^{-3}$
- ▶ $T_e = T_i = 1000 \text{ K} = 0.0862 \text{ eV}$
- ▶ $m_i = \cancel{16 \text{ amu}} \rightarrow 1/16 \text{ amu} \gg m_e$ (speed up sim.)
- ▶ $u = \cancel{7000 \text{ m/s}} \rightarrow 112000 \text{ m/s}$
(maintain correct wake dynamics)

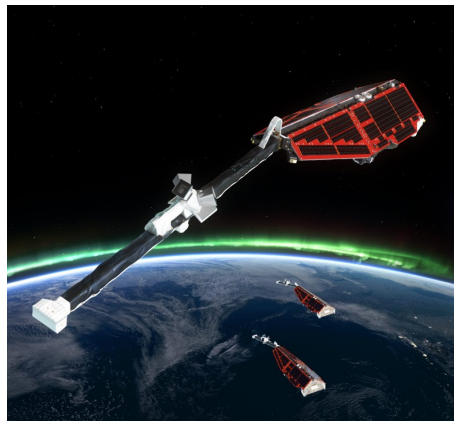


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Step 1:
Mesh generation
(25.01.22)

sphere_0.5R.geo

Gmsh

sphere_0.5R.msh

msh2topo

Step 2:
Simulation
(27.01.22)

pictetra.dat

PTetra

meshpic.dat →
sphere_0.5R.topo

alujlu.prcnd

**.vtk*

pictetra.hst

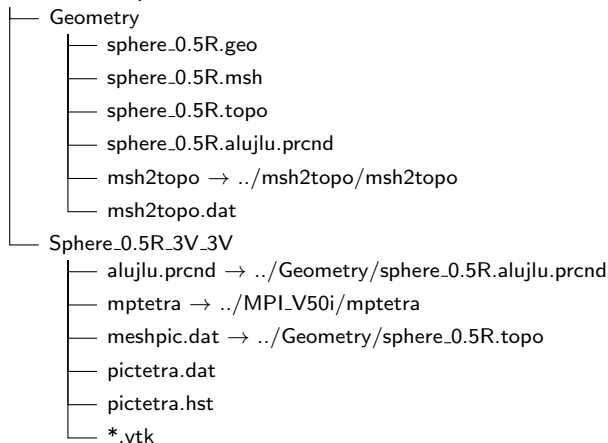
Step 3:
Post-processing
(03.02.22)

ParaView

Python, etc.

Naming conventions

PTetraWorkshop



Geometry names:

sphere_0.5R.geo
 shape radius in Debye lengths

cylinder_0.2R_10L.geo
 length in Debye lengths

Simulation folder names:

Sphere_0.5R_1V_3V_112kms/
 same as geometry voltages of drift speed
 the two halves

Sphere_0.5R_FV_112kms/
 floating voltage